#### The scenario of perfect fluid hydrodynamics

### in A+A collisions at RHIC T. Csörgő MTA KFKI RMKI, Budapest, Hungary

#### •Introduction:

- RHIC Scientists serve up "Perfect Liquid", April 18, 2005
- **BRAHMS, PHENIX, PHOBOS, STAR White Papers, NPA, 2005**
- AIP Top Physics Story 2005
- Hydrodynamics and scaling of soft observables
  - Exact (i.e. not numerical) integrals of fluid dynamics
    - non-relativistic and relativistic solutions
  - Evidence for hydrodynamic scaling in RHIC data
    - Scaling of slope parameters
    - Scaling of Bose-Einstein / HBT radii
    - Universal scaling of elliptic and higher order flows
- Intermediate p<sub>t</sub> region: breaking of the hydro scaling
  - 1

### **Location of experiments: BNL**

- RHIC:  $Au + Au @ E_{cms} = 200 AGeV$ 
  - Au+Au, Cu+Cu, pp+pp, d+Au collisions
  - 4 experiments: **BRAHMS, PHENIX, PHOBOS, STAR**
- **Brazilian participation: PHENIX and STAR:**



#### PHENIX

- photons, electrons muons and hadrons
- Investigates all stages of the reaction
- Penetrating probes: early state of the reaction
- Hadrons



#### 1<sup>st</sup> milestone: new phenomena



Suppression of high pt particle production in Au+Au collisions at RHIC

4

### 2<sup>nd</sup> milestone: new form of matter



## 3<sup>rd</sup> milestone: Top Physics Story 2005

Cím 🕘 http://www.aip.org/pnu/2005/split/757-1.html	
AMERICAN IN	STITUTE OF PHYSICS
Physics News Update The AIP Bulletin of Physics News	
Number 757 #1, December 7, 2005 by Phil Schewe and Ben Stein	
Subscribe to Physics News	The Top Physics Stories for 2005
Update Physics News Graphics	At the Relativistic Heavy Ion Collider (RHIC) on Long Island, the four large detector groups agreed, for the first time, on a consensus interpretation of several year's worth of high-energy ion collisions: the fireball made in these collisions a sort of stand-in for the primordial
<u>Physical</u> <u>Review Focus</u>	universe only a few microseconds after the big bang was not a gas of weakly interacting quarks and gluons as earlier expected, but something more like a liquid of strongly interacting quarks and gluons ( <u>PNU 728</u> ).
Physics News Links	Other top physics stories for 2005 include, in general chronological order of their appearance throughout the year, the following:
Archives 2006	the arrival of the Cassini spacecraft at Saturn and the successful landing of the Huygens probe on the moon Titan ( <u>PNU 716</u> );
<u>2005</u> 2004	the development of lasing in silicon ( <u>Nature 17 February</u> );

#### http://arxiv.org/abs/nucl-ex/0410003

### **Discovering New Laws**

"In general we look for a new law by the following process.

First we guess it

Then we compare the consequences of the guess to see

what would be implied if this law that we guessed is right.

Then we compare the result of the computation to nature,

with experiment or experience, compare it directly with observation,

<u>to see if it works.</u>

<u>If it disagrees with experiment it is wrong.</u>

In that simple statement is the key to science. It does not make any difference how beautiful your guess is. It does not make any difference how smart you are, who made the guess, or what his name is if it disagrees with experiment it is wrong."

#### <u>/R.P. Feynman/</u>

### **An observation:**



Inverse slopes T of single particle  $p_t$  distribution increase linearly with mass:  $T = T_0 + m < u_t > 2$ 

Increase is stronger in more head-on collisions. Suggests collective radial flow, local thermalization and hydrodynamics

T. Csörgő @ IFT Sao Paulo, 2006/8/23

8

### **Notation for fluid dynamics**

#### • nonrelativistic hydro:

- t: time,
- r: coordinate 3-vector,  $r = (r_x, r_y, r_z)$ ,
- m: mass,
- field i.e. (t,r) dependent variables:
  - n: number density,
  - p: pressure,
  - ε: energy density,
  - T: temperature,
  - v: velocity 3-vector,  $v = (v_x, v_y, v_z)$ ,
- relativistic hydro:
  - $x^{\mu}$ : coordinate 4-vector,  $x^{\mu} = (t, r_x, r_y, r_z)$ ,
  - $k^{\mu}$ : momentum 4-vector,  $k^{\mu}$  = (E,  $k_x$ ,  $k_y$ ,  $k_z$ ),  $k^{\mu} k_{\mu}$  = m<sup>2</sup>,
- additional fields in relativistic hydro:

 $u^{\mu}$  : velocity 4-vector,  $u^{\mu} = \gamma (1, v_x, v_y, v_z), \quad u^{\mu} u_{\mu} = 1,$  $g^{\mu\nu}$ : metric tensor,  $g^{\mu\nu} = diag(1, -1, -1, -1),$ 

 $T^{\mu\nu}$ : energy-momentum tensor.

## **Nonrelativistic dynamics of perfect fluids**

#### • Equations of nonrelativistic hydro:

local conservation of

charge: continuity

momentum: Euler

energy

$$\partial_t n + \nabla \cdot (n\mathbf{v}) = 0,$$
  

$$\partial_t \mathbf{v} + (\mathbf{v} \cdot \nabla)\mathbf{v} = -(\nabla p)/(mn),$$
  

$$\partial_t \epsilon + \nabla \cdot (\epsilon \mathbf{v}) = -p\nabla \cdot \mathbf{v},$$

• Not closed, EoS needed:

$$p = nT$$
,  $\epsilon = \kappa(T)nT$ ,

 Perfect fluid: definitions are equivalent, term used by PDG # 1: no bulk and shear viscosities, and no heat conduction. # 2: T<sup>µv</sup> = diag(e,-p,-p,-p) in the local rest frame.

#### ideal fluid: ambiguously defined term, discouraged

#1: keeps its volume, but conforms to the outline of its container#2: an inviscid fluid

10

## **Input from lattice: EoS of QCD Matter**

#### <u>Old</u> idea: Quark Gluon Plasma "Ionize" nucleons with heat "Compress" them with density New state(s?) of matter

Z. Fodor and S.D. Katz: critical end point of 1st order phase tr even at finite baryon density, cross over like transition. (hep-lat/0106002, hep-lat/0402006) T<sub>c</sub>=176±3 MeV (~2 terakelvin) (hep-ph/0511166)





General input for hydro:  $p(\mu,T)$ LQCD for RHIC region:  $p \sim p(T)$ ,  $c_s^2 = \delta p / \delta e = c_s^2(T) = 1/\kappa(T)$ 

It's in the family analytic hydro solutions!

## New parametric, ellipsoidal hydro solutions

# Ansatz: the density n (and T and $\epsilon$ ) depend on coordinates only through a scale parameter s

• T. Cs. Acta Phys. Polonica B37 (2006) 1001, hep-ph/0111139

$$n = f(t)g(s).$$
  
$$\partial_t n = f'(t)g(s) + f(t)g'(s)\partial_t s,$$
  
$$\nabla(vn) = f(t)g(s)\nabla v + f(t)g'(s)v\nabla s.$$

Scale parameters: (X,Y,Z) = (X(t), Y(t), Z(t))

 $s = \frac{r_x^2}{V^2} + \frac{r_y^2}{V^2} + \frac{r_z^2}{Z^2}$ 

$$f(t) = \frac{X_0 Y_0 Z_0}{XYZ}$$

$$\frac{f'(t)}{f(t)} = -\nabla v,$$

$$\partial_t s + v\nabla s = 0$$

$$v = \left(\frac{\dot{X}}{X} r_x, \frac{\dot{Y}}{Y} r_y, \frac{\dot{Z}}{Z} r_z\right)$$

Density=const on ellipsoids. Directional Hubble flow. g(s): arbitrary scaling function. Notation:  $n \sim v(s)$ ,  $T \sim \tau(s)$  etc.

### New exact, ellipsoidal hydro solutions

#### A new family of PARAMETRIC, exact, scale-invariant solutions

T. Cs. Acta Phys. Polonica B37 (2006) 1001, hep-ph/0111139

Volume is introduced as V = XYZ

$$n(t, \mathbf{r}) = n_0 \frac{V_0}{V} \nu(s)$$
  

$$\mathbf{v}(t, \mathbf{r}) = \left(\frac{\dot{X}}{X} r_x, \frac{\dot{Y}}{Y} r_y, \frac{\dot{Z}}{Z} r_z\right)$$
  

$$T(t, \mathbf{r}) = T_0 \left(\frac{V_0}{V}\right)^{1/\kappa} \mathcal{T}(s)$$
  

$$\nu(s) = \frac{1}{\mathcal{T}(s)} \exp\left(-\frac{T_i}{2T_0} \int_0^s \frac{du}{\mathcal{T}(u)}\right)$$

$$s = \frac{r_x^2}{X^2} + \frac{r_y^2}{Y^2} + \frac{r_z^2}{Z^2}$$

For  $\kappa = \kappa(T)$  exact solutions, see T. Cs, S.V. Akkelin, Y. Hama, B. Lukács, Yu. Sinyukov, hep-ph/0108067, Phys.Rev.C67:034904,2003

The dynamics is reduced to coupled, nonlinear but ordinary differential equations for the scales X,Y,Z

$$X\ddot{X} = Y\ddot{Y} = Z\ddot{Z} = \frac{T_i}{m} \left(\frac{V_0}{V}\right)^{1/\kappa}$$

All hydro problems (initial conditions, role of EoS, freeze-out conditions) can be easily illustrated and understood on the equivalent problem:

a classical potential motion of a mass-point in a conservative potential (a shot)!

Note: temperature scaling function  $\tau(s)$  remains arbitrary!  $\nu(s)$  depends on  $\tau(s)$ . -> FAMILY of solutions.

## From fluid expansion to potential motion

#### **Dynamics of pricipal axis:**



The role of initial boundary conditions, EoS and freeze-out in hydro can be understood from potential motion!

## **Initial boundary conditions**

#### From the new family of exact solutions, the initial conditions:

#### **Initial coordinates:**

(nuclear geometry + time of thermalization)

**Initial velocities:** 



 $T_0$ 

 $n_0$ 

 $\tau(s)$ 

 $X_0 Y_0 Z_0$ 

(pre-equilibrium+ time of thermalization)

**Initial temperature:** 

**Initial density:** 

#### **Initial profile function:**

(energy deposition and pre-equilibrium process)



### **Role of initial temperature profile**

- Initial temperature profile = arbitrary positive function
- Infinitly rich class of solutions
- Matching initial conditions for the density profile
  - T. Cs. Acta Phys. Polonica B37 (2006) 1001, hep-ph/0111139

$$\nu(s) = \frac{1}{\mathcal{T}(s)} \exp\left(-\frac{T_i}{2T_0} \int_0^s \frac{du}{\mathcal{T}(u)}\right)$$

Homogeneous temperature ⇒ Gaussian density

$$\nu(s) = \exp(-s/2), \quad \mathcal{T}(s) = 1.$$

$$s = \frac{r_x^2}{X^2} + \frac{r_y^2}{Y^2} + \frac{r_z^2}{Z^2}$$

Buda-Lund profile:

$$\mathcal{T}(s) = \frac{1}{1+bs}$$
  
$$\nu(s) = (1+bs) \exp\left[-\frac{T_i}{2T_0}(s+bs^2/2)\right]$$

#### Zimányi-Bondorf-Garpman profile:

$$\mathcal{T}(s) = (1-s)\Theta(1-s)$$
  
$$\nu(s) = (1-s)^{\alpha}\Theta(1-s)$$

## **Illustrated initial T-> density profiles**



## Final (freeze-out) boundary conditions

 $T_{f}$ 

 $n_f$ 

(s)

From the new exact hydro solutions, the conditions to stop the evolution:

Freeze-out temperature:

**Final coordinates:** 



(cancel from measurables, diverge)

**Final velocities:** 

(determine observables, tend to constants)

#### **Final density:**

(cancels from measurables, tends to 0)

#### **Final profile function:**

(= initial profile function! from solution)



### **Role of the Equation of States:**

#### The potential depends

**on**  $\kappa = \delta \varepsilon / \delta p$ :





Time evolution of the scales (X,Y,Z) follows a classic potential motion. Scales at freeze out determine the observables. Info on history <u>LOST</u>! No go theorem - constraints on initial conditions (information on spectra, elliptic flow of penetrating probels) indispensable.

The arrow hits the target, but can one determine g from this information??

19

### **Initial conditions <-> Freeze-out conditions:**

#### Different initial conditions

but

same freeze-out conditions

ambiguity!

Penetrating probes radiate through the time evolution!



### **Illustrations of exact hydro results**

Propagate the hydro solution in time numerically:



21

 $R_{x}(t), R_{y}(t), R_{z}(t)$ 

## Solution of the "HBT puzzle"



Geometrical sizes keep on increasing. Expansion velocities tend to constants. HBT radii  $R_x$ ,  $R_y$ ,  $R_z$  approach a direction independent constant. Slope parameters tend to direction dependent constants. General property, independent of initial conditions - a beautiful exact result.

### **Geometrical & thermal & HBT radii**

23



Geometrical radii Thermal radii HBT radii **3d analytic hydro: exact time evolution** 

geometrical size (fugacity ~ const) Thermal sizes (velocity ~ const) HBT sizes (phase-space density ~ const)

HBT dominated by the smaller of the geometrical and thermal scales

nucl-th/9408022, hep-ph/9409327 hep-ph/9509213, hep-ph/9503494

HBT radii approach a constant of time HBT volume becomes spherical HBT radii -> thermal ~ constant sizes

> hep-ph/0108067, nucl-th/0206051 animation by Máté Csanád

## **Scaling predictions of fluid dynamics**

$$T'_x = T_f + m \dot{X}_f^2 ,$$
  

$$T'_y = T_f + m \dot{Y}_f^2 ,$$
  

$$T'_z = T_f + m \dot{Z}_f^2 .$$

- Slope parameters increase linearly with mass
- Elliptic flow is a universal function its variable w is proportional to transverse kinetic energy and depends on slope differences.

$$v_2 = \frac{I_1(w)}{I_0(w)}$$

$$w = \frac{k_t^2}{4m} \left(\frac{1}{T_y'} - \right)$$

$$w = \frac{E_K}{2T_*}\varepsilon$$

Inverse of the HBT radii increase linearly with mass analysis shows that they are asymptotically the same

Relativistic correction:  $m \rightarrow m_t$ 

hep-ph/0108067, nucl-th/0206051

$$R'_{x}^{-2} = X_{f}^{-2} \left( 1 + \frac{m}{T_{f}} \dot{X}_{f}^{2} \right),$$
$$R'_{y}^{-2} = Y_{f}^{-2} \left( 1 + \frac{m}{T_{f}} \dot{Y}_{f}^{2} \right),$$
$$R'_{z}^{-2} = Z_{f}^{-2} \left( 1 + \frac{m}{T_{f}} \dot{Z}_{f}^{2} \right).$$

#### **Relativistic Perfect Fluids**

#### **Rel. hydrodynamics of perfect fluids is defined by:**

$$\partial_{\mu} \left( n u^{\mu} \right) = 0$$
$$\partial_{\mu} T^{\mu\nu} = 0$$

T

$$T^{\mu\nu} = (\varepsilon + p) u^{\mu} u^{\nu} - p g^{\mu\nu}$$

T. Csörgő @ IFT Sao Paulo, 2006/8/23

A recent family of exact solutions: nucl-th/0306004

$$u^{\mu} = \frac{x^{\mu}}{\tau}$$

$$n(t, \mathbf{r}) = n_0 \left(\frac{\tau_0}{\tau}\right)^3 \mathcal{V}(s)$$

$$u^{\mu} \partial_{\mu} p + (\epsilon + p) u^{\mu} \partial_{\mu} u_{\nu} - \partial_{\nu} p = 0,$$

$$u^{\mu} \partial_{\mu} T + \frac{1}{\kappa} T \partial_{\mu} u^{\mu} = 0.$$

$$r(t, \mathbf{r}) = p_0 \left(\frac{\tau_0}{\tau}\right)^{3+3/\kappa} \frac{1}{\mathcal{V}(s)}$$

$$s = \frac{r_x^2}{X^2} + \frac{r_y^2}{Y^2} + \frac{r_z^2}{Z^2}, \quad \epsilon = mn + \kappa p,$$

$$p = nT.$$

**Overcomes two shortcomings of Bjorken's solution:** Yields finite rapidity distribution, includes transverse flow  $u^{\mu}\partial_{\mu}u_{\nu}=0$ **Hubble flow**  $\Rightarrow$  **lack of acceleration**: Accelerating, new rel. hydro solutions: nucl-th/0605070

25

### **Solutions of Relativistic Perfect Fluids**

#### A new family of exact solutions:

T. Cs, M. I. Nagy, M. Csanád: nucl-th/0605070

#### **Overcomes two shortcomings of Bjorken's solution:**

Finite Rapidity distribution ~ Landau's solution

**Includes relativistic acceleration** 

in 1+1 and 1+3 spherically symmetric



26

#### **Animation of the new exact solution**



r

## nucl-th/0605070: advanced estimate of $\varepsilon_0$

Width of dn/dy distribution is due to acceleration: acceleration yields longitudinal explosion, thus Bjorken estimate underestimates initial energy density by 50 %:



$$\epsilon_{0} = \frac{\left\langle m_{t} \right\rangle}{R^{2} \pi \tau_{0}} \frac{dn}{d\eta_{0}} = \epsilon_{Bj} \frac{dy}{d\eta_{f}} \frac{d\eta_{f}}{d\eta_{0}}$$

$$\frac{\varepsilon_0}{\varepsilon_{Bj}} = \frac{\alpha}{\alpha - 2} \left(\frac{\tau_f}{\tau_0}\right)^{\frac{1}{\alpha - 2}} = \left(2\lambda - 1\right) \left(\frac{\tau_f}{\tau_0}\right)^{\lambda - 1}$$

### nucl-th/0605070: advanced estimate of $\varepsilon_0$

M. Csanád-> fits to BRAHMS dn/d $\eta$  data dn/d $\eta$  widths yields correction factors of ~ 2.0 - 2.2 Yields inital energy density of  $\epsilon_0 \sim 10-30 \text{ GeV/fm}^3$ 

a correction of  $\mathcal{E}_0/\mathcal{E}_{Bj} \sim 2$  as compared to PHENIX White Paper!



#### T. Csörgő @ IFT Sao Paulo, 2006/8/23

29

## **Principles for Buda-Lund hydro model**

- Analytic expressions for all the observables
- 3d expansion, local thermal equilibrium, symmetry
- Goes back to known exact hydro solutions:
  - nonrel, Bjorken, and Hubble limits, 1+3 d ellipsoids
  - but phenomenology, extrapolation for unsolved cases
- Separation of the Core and the Halo
  - Core: perfect fluid dynamical evolution
  - Halo: decay products of long-lived resonances
- Missing links: phenomenology needed
  - search for accelerating ellipsoidal rel. solutions
  - first accelerating rel. solution: nucl-th/0605070





## A useful analogy

#### Fireball at RHIC ⇔ our Sun



31

### **Buda-Lund hydro model**

#### The general form of the emission function:

$$S_c(x,p)d^4x = \frac{g}{(2\pi)^3} \frac{p^{\mu}d^4\Sigma_{\mu}(x)}{\exp\left(\frac{p^{\nu}u_{\nu}(x)}{T(x)} - \frac{\mu(x)}{T(x)}\right) + s_q}$$

#### **Calculation of observables with core-halo correction:**

$$N_1(p) = \frac{1}{\sqrt{\lambda_*}} \int d^4x S_c(p, x)$$

$$C(Q,p) = 1 + \left|\frac{\tilde{S}(Q,p)}{\tilde{S}(0,p)}\right|^2 = 1 + \lambda_* \left|\frac{\tilde{S}_c(Q,p)}{\tilde{S}_c(0,p)}\right|^2$$

#### **Assuming profiles for** flux, temperature, chemical potential and flow T. Csörgő @ IFT Sao Paulo, 2006/8/23

## **The generalized Buda-Lund model**

The original model was for axial symmetry only, central coll. In its general hydrodynamical form:

Based on 3d relativistic and non-rel solutions of perfect fluid dynamics:

$$S_c(x,p)d^4x = \frac{g}{(2\pi)^3} \frac{p^{\mu} d^4 \Sigma_{\mu}(x)}{\exp\left(\frac{p^{\nu} u_{\nu}(x)}{T(x)} - \frac{\mu(x)}{T(x)}\right) + s_q}$$

Have to assume special shapes:

**Generalized Cooper-Frye prefactor:** 

$$p^{\mu}d^{4}\Sigma_{\mu}(x) = p^{\mu}u_{\mu}(x)H(\tau)d^{4}x$$

$$H(\tau) = \frac{1}{(2\pi\Delta\tau^2)^{1/2}} \exp\left(-\frac{(\tau - \tau_0)^2}{2\Delta\tau^2}\right)$$

**Four-velocity distribution:** 

$$\begin{aligned} u^{\mu} &= (\gamma, \sinh \eta_x, \sinh \eta_y, \sinh \eta_z) \\ \frac{1}{T(x)} &= \frac{1}{T_0} \left( 1 + \frac{T_0 - T_s}{T_s} s \right) \left( 1 + \frac{T_0 - T_e}{T_e} \frac{(\tau - \tau_0)^2}{2\Delta \tau^2} \right) \\ \frac{\mu(x)}{T(x)} &= \frac{\mu_0}{T_0} - s \end{aligned}$$

$$s &= \frac{r_x^2}{X^2} + \frac{r_y^2}{Y^2} + \frac{r_z^2}{Z^2} \end{aligned}$$

**Temperature:** 

**Fugacity:** 

#### T. Csörgő @ IFT Sao Paulo, 2006/8/23

33

## Scaling predictions of non-rel hydro

$$T'_x = T_f + m \dot{X}_f^2 ,$$
  

$$T'_y = T_f + m \dot{Y}_f^2 ,$$
  

$$T'_z = T_f + m \dot{Z}_f^2 .$$

- Slope parameters increase linearly with mass
- Elliptic flow is a <u>universal function</u> and variable w is proportional to transverse kinetic energy and depends on slope differences.

$$v_2 = \frac{I_1(w)}{I_0(w)}$$

$$w = \frac{k_t^2}{4m} \left(\frac{1}{T_y'} - \right)$$

$$w = \frac{E_K}{2T_*}\varepsilon$$

Inverse of the HBT radii increase linearly with mass analysis shows that they are asymptotically the same

Relativistic correction:  $m \rightarrow m_t$ 

hep-ph/0108067, nucl-th/0206051

$$R'_{x}^{-2} = X_{f}^{-2} \left( 1 + \frac{m}{T_{f}} \dot{X}_{f}^{2} \right),$$
$$R'_{y}^{-2} = Y_{f}^{-2} \left( 1 + \frac{m}{T_{f}} \dot{Y}_{f}^{2} \right),$$
$$R'_{z}^{-2} = Z_{f}^{-2} \left( 1 + \frac{m}{T_{f}} \dot{Z}_{f}^{2} \right).$$

# Scaling predictions: Buda-Lund hydro

$$T_x = T_0 + \overline{m}_t \dot{X}^2 \frac{T_0}{T_0 + \overline{m}_t a^2},$$

$$\overline{m}_t = m_t \cosh(\eta_s - y).$$

- Slope parameters increase linearly with transverse mass
- Elliptic flow is same universal function.
- Scaling variable w is prop. to generalized transv. kinetic energy and depends on effective slope diffs.

$$v_2 = \frac{I_1(w)}{I_0(w)} \qquad w = \frac{E_F}{2T}$$

$$w = \frac{E_K}{2T_*}\varepsilon$$

$$\varepsilon = \frac{T_x - T_y}{T_x + T_y}.$$

$$E_K = \frac{p_t^2}{2\overline{m}_t} \qquad \qquad \frac{1}{T_*} = \frac{1}{2} \left( \frac{1}{T_x} + \frac{1}{T_y} \right).$$

Inverse of the HBT radii increase linearly with mass analysis shows that they are asymptotically the same

Relativistic correction:  $m \rightarrow m_{t}$ 

hep-ph/0108067, nucl-th/0206051

$$\frac{1}{R_{i,i}^2} = \frac{B(x_s, p)}{B(x_s, p) + s_q} \left(\frac{1}{X_i^2} + \frac{1}{R_{T,i}^2}\right)$$

$$\frac{1}{R_{T,i}^2} = \frac{m_t}{T_0} \left( \frac{a^2}{X_i^2} + \frac{\dot{X}_i^2}{X_i^2} \right)$$

### **Some analytic Buda-Lund results**

#### **HBT radii widths:**

$$\frac{1}{R_{i,i}^2} = \frac{B(x_s, p)}{B(x_s, p) + s_q} \left(\frac{1}{X_i^2} + \frac{1}{R_{T,i}^2}\right) = \frac{1}{R_{T,i}^2} = \frac{m_t}{T_0} \left(\frac{a^2}{X_i^2} + \frac{\dot{X}_i^2}{X_i^2}\right)$$

$$a^{2} = \frac{T_{0} - T_{s}}{T_{s}} = \left\langle \frac{\Delta T}{T} \right\rangle_{r}$$

$$\overline{m}_t = m_t \cosh(\eta_s - y).$$

$$\frac{1}{T_*} = \frac{1}{2} \left( \frac{1}{T_x} + \frac{1}{T_y} \right).$$

**Slopes, effective temperatures:** 

$$T_x = T_0 + \overline{m}_t \dot{X}^2 \frac{T_0}{T_0 + \overline{m}_t a^2},$$

#### Flow coefficients are <u>universal</u>:

$$v_{2n} = \frac{I_n(w)}{I_0(w)} \qquad w = \frac{E_K}{2T_*}\varepsilon$$
$$v_{2n+1} = 0$$

$$\varepsilon = \frac{T_x - T_y}{T_x + T_y}.$$

$$E_K = \frac{p_t^2}{2\overline{m}_t}$$

## Buda-Lund hydro and Au+Au@RHIC



37

## Femptoscopy signal of sudden hadronization



38

### Confirmation



see nucl-th/0310040 and nucl-th/0403074, R. Lacey@QM2005/ISMD 2005 A. Ster @ QM2005.

## Hydro scaling of slope parameters



## Hydro scaling of Bose-Einstein/HBT radii



same slopes ~ fully developed, 3d Hubble flow

## Hydro scaling of elliptic flow



G. Veres, PHOBOS data, proc QM2005 Nucl. Phys. A774 (2006) in press

# Hydro scaling of $v_2$ and $\sqrt{s}$ dependence

# PHOBOS, nucl-ex/0406021



#### T. Csörgő @ IFT Sao Paulo, 2006/8/23

43

### **Universal scaling and v<sub>2</sub>(centrality,η)**

# PHOBOS, nucl-ex/0407012



### **Universal v2 scaling and PID dependence**

# **PHENIX**, nucl-ex/0305013



T. Csörgő @ IFT Sao Paulo, 2006/8/23

45

### Universal scaling and fine structure of v2

# STAR, nucl-ex/0409033



# Universal v<sub>2</sub> scaling predicted in 2003



## Summary on universal hydro scaling of v<sub>2</sub>

48



Black line: Theoretically predicted, universal scaling function from analytic works on perfect fluid hydrodynamics:



hep-ph/0108067, nucl-th/0310040

# Scaling and scaling violations



# **Understanding hydro results**



# **Summary**

Au+Au elliptic flow data at RHIC satisfy the UNIVERSAL scaling laws predicted (2001, 2003) by the (Buda-Lund) hydro model, based on exact solutions of PERFECF FLUID hydrodynamics

quantitative evidence for a perfect fluid in Au+Au at RHIC

scaling breaks in p<sub>t</sub> at ~ 1.5 GeV, in rapidity at ~|y| >  $y_{may}$  - 0.5 Search for establishing the domain of applicability started. Scaling of HBT radii and spectra: first tests passed, further tests going on.