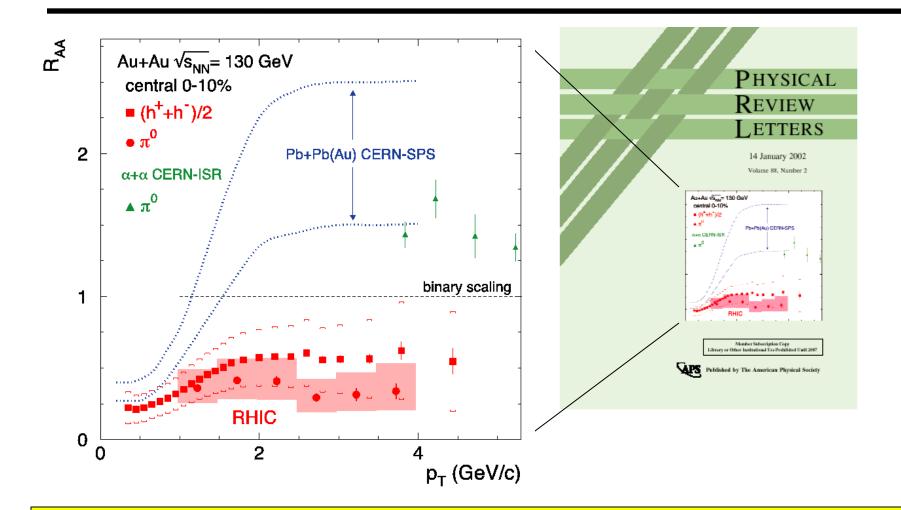
## The perfect fluid

# in A+A collisions at RHIC T. Csörgő MTA KFKI RMKI, Budapest, Hungary

#### •Introduction:

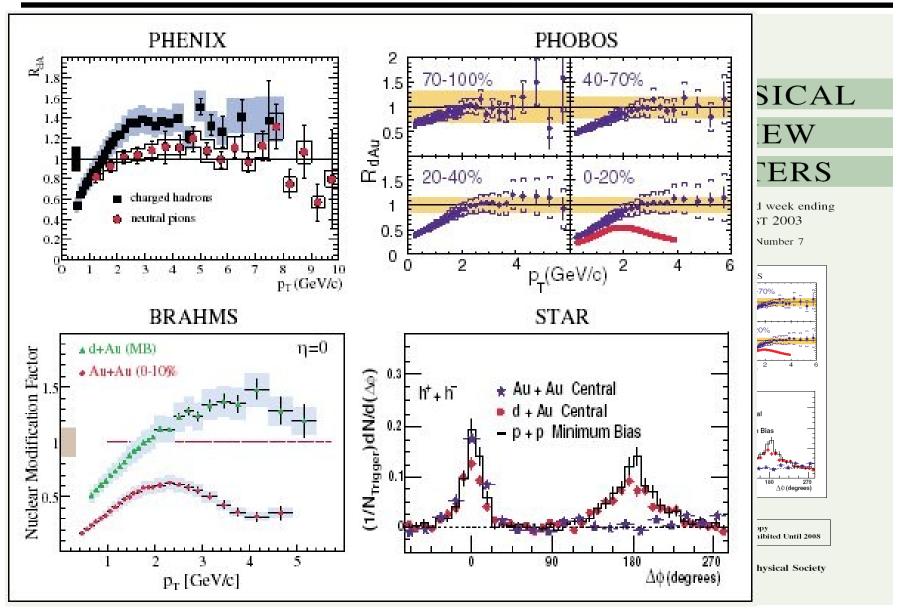
- RHIC Scientists serve up "Perfect Liquid", April 18, 2005
- **□** BRAHMS, PHENIX, PHOBOS, STAR White Papers, NPA, 2005
- **□ AIP Top Physics Story 2005**
- Hydrodynamics and scaling of soft observables
  - Exact (i.e. not numerical) integrals of fluid dynamics
    - non-relativistic and relativistic solutions
  - Evidence for hydrodynamic scaling in RHIC data
    - Scaling of slope parameters
    - Scaling of Bose-Einstein / HBT radii
    - Universal scaling of elliptic and higher order flows
- Intermediate p<sub>t</sub> region: breaking of the hydro scaling

## 1st milestone: new phenomena

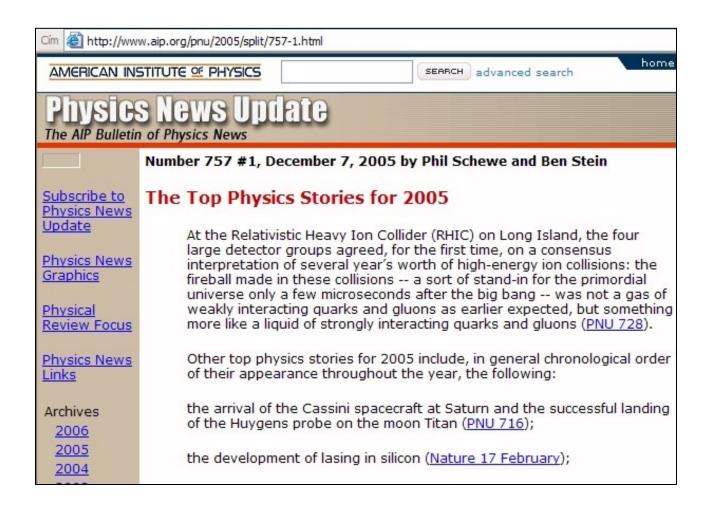


Suppression of high p<sub>t</sub> particle production in Au+Au collisions at RHIC

## 2<sup>nd</sup> milestone: new form of matter



# 3<sup>rd</sup> milestone: Top Physics Story 2005



http://arxiv.org/abs/nucl-ex/0410003

## **Discovering New Laws**

"In general we look for a new law by the following process.

First we guess it.

Then we compare the consequences of the guess to see what would be implied if this law that we guessed is right.

Then we compare the result of the computation to nature, with experiment or experience, compare it directly with observation, to see if it works.

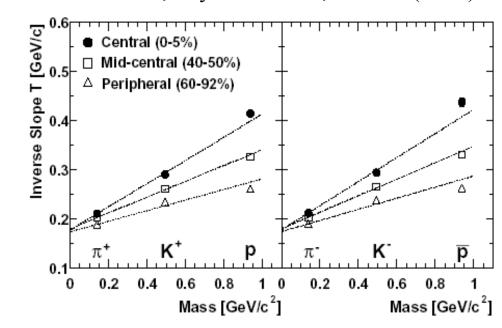
If it disagrees with experiment it is wrong.

In that simple statement is the key to science.
It does not make any difference how beautiful your guess is.
It does not make any difference how smart you are,
who made the guess, or what his name is —
if it disagrees with experiment it is wrong."

/R.P. Feynman/

## An observation:

PHENIX, Phys. Rev. C69, 034909 (2004)



<u>Inverse slopes T of single particle p<sub>t</sub> distribution</u> increase linearly with mass:

$$T = T_0 + m < u_{\underline{t}} > 2$$

Increase is stronger in more head-on collisions.
Suggests collective radial flow, local thermalization and hydrodynamics

## **Notation for fluid dynamics**

```
nonrelativistic hydro:
       time,
    r: coordinate 3-vector, r = (r_x, r_y, r_z),
    m: mass,

    field i.e. (t,r) dependent variables:

    n: number density,
    p: pressure,
    ε: energy density,
    T: temperature,
    v: velocity 3-vector, v = (v_x, v_y, v_z),
relativistic hydro:
    x^{\mu}: coordinate 4-vector, x^{\mu} = (t, r_x, r_y, r_z),
    k^{\mu}: momentum 4-vector, k^{\mu} = (E, k_x, k_y, k_z), k^{\mu} k_{\mu} = m^2,

    additional fields in relativistic hydro:

    u^{\mu}: velocity 4-vector, u^{\mu} = \gamma (1, v_x, v_y, v_z), \qquad u^{\mu} u_{\mu} = 1,
    g^{\mu\nu}: metric tensor, g^{\mu\nu} = diag(1,-1,-1,-1),
    T<sup>μν</sup>: energy-momentum tensor.
```

## Nonrelativistic dynamics of perfect fluids

- Equations of nonrelativistic hydro:
  - local conservation of

**charge: continuity** 

momentum: Euler

energy

$$\partial_t n + \nabla \cdot (n\mathbf{v}) = 0,$$
  

$$\partial_t \mathbf{v} + (\mathbf{v} \cdot \nabla)\mathbf{v} = -(\nabla p)/(mn),$$
  

$$\partial_t \epsilon + \nabla \cdot (\epsilon \mathbf{v}) = -p\nabla \cdot \mathbf{v},$$

Not closed, EoS needed:

$$p = nT$$
,  $\epsilon = \kappa(T)nT$ ,

Perfect fluid: definitions are equivalent, term used by PDG

# 1: no bulk and shear viscosities, and no heat conduction.

# 2:  $T^{\mu\nu}$  = diag(e,-p,-p,-p) in the local rest frame.

ideal fluid: ambiguously defined term, discouraged

#1: keeps its volume, but conforms to the outline of its container

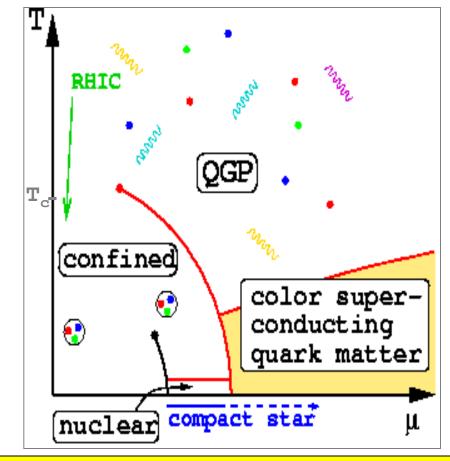
#2: an inviscid fluid

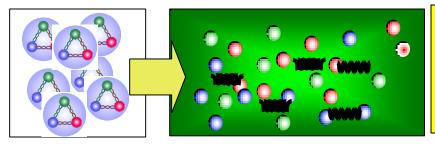
## Input from lattice: EoS of QCD Matter

Old idea: Quark Gluon Plasma
"Ionize" nucleons with heat
"Compress" them with density
New state(s?) of matter

Z. Fodor and S.D. Katz:
critical end point of 1st order phase tr
even at finite baryon density,
cross over like transition.
(hep-lat/0106002, hep-lat/0402006)

 $T_c=176\pm3$  MeV (~2 terakelvin) (hep-ph/0511166)





General input for hydro:  $p(\mu,T)$ LQCD for RHIC region:  $p \sim p(T)$ ,  $c_s^2 = \delta p/\delta e = c_s^2(T) = 1/\kappa(T)$ It's in the family analytic hydro solutions!

## New exact, parametric hydro solutions

## Ansatz: the density n (and T and $\varepsilon$ ) depend on coordinates only through a scale parameter s

T. Cs. Acta Phys. Polonica B37 (2006) 1001, hep-ph/0111139

$$n = f(t)g(s).$$

$$\partial_t n = f'(t)g(s) + f(t)g'(s)\partial_t s,$$

$$\nabla(vn) = f(t)g(s)\nabla v + f(t)g'(s)v\nabla s.$$

**Principal axis of ellipsoid:** (X,Y,Z) = (X(t), Y(t), Z(t))

$$f(t) = \frac{X_0 Y_0 Z_0}{XYZ}$$

$$s = \frac{r_x^2}{X^2} + \frac{r_y^2}{Y^2} + \frac{r_z^2}{Z^2}$$

**Density=const on ellipsoids.** g(s): arbitrary scaling function.

$$\frac{f'(t)}{f(t)} = -\nabla v, \quad \partial_t s + v \nabla s = 0$$

$$\partial_t s + v \nabla s = 0$$

$$v = \left(\frac{\dot{X}}{X} r_x, \frac{\dot{Y}}{Y} r_y, \frac{\dot{Z}}{Z} r_z\right)$$

**Directional Hubble flow.** 

Notation:  $n \sim v(s)$ ,  $T \sim \tau(s)$  etc.

## New exact, ellipsoidal hydro solutions

#### A new family of PARAMETRIC, exact, scale-invariant solutions

T. Cs. Acta Phys. Polonica B37 (2006) 1001, hep-ph/0111139

Volume is introduced as V = XYZ

$$n(t, \mathbf{r}) = n_0 \frac{V_0}{V} \nu(s)$$

$$\mathbf{v}(t, \mathbf{r}) = \left(\frac{\dot{X}}{X} r_x, \frac{\dot{Y}}{Y} r_y, \frac{\dot{Z}}{Z} r_z\right)$$

$$T(t, \mathbf{r}) = T_0 \left(\frac{V_0}{V}\right)^{1/\kappa} \mathcal{T}(s)$$

$$\nu(s) = \frac{1}{\mathcal{T}(s)} \exp\left(-\frac{T_i}{2T_0} \int_0^s \frac{du}{\mathcal{T}(u)}\right)$$

$$s = \frac{r_x^2}{X^2} + \frac{r_y^2}{Y^2} + \frac{r_z^2}{Z^2}$$

For  $\kappa = \kappa(T)$  exact solutions, see

T. Cs, S.V. Akkelin, Y. Hama,

B. Lukács, Yu. Sinyukov,

hep-ph/0108067, Phys.Rev.C67:034904,2003

The dynamics is reduced to coupled, nonlinear but ordinary differential equations for the scales X,Y,Z

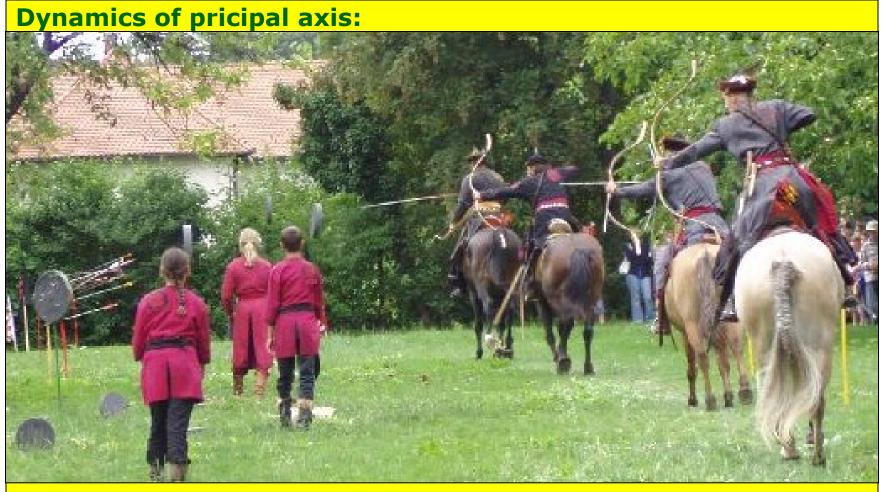
$$X\ddot{X} = Y\ddot{Y} = Z\ddot{Z} = \frac{T_i}{m} \left(\frac{V_0}{V}\right)^{1/\kappa}$$

All hydro problems (initial conditions, role of EoS, freeze-out conditions) can be <u>easily</u> illustrated and understood on the <u>equivalent problem</u>:

a classical potential motion of a mass-point in a conservative potential (a shot)!

Note: temperature scaling function  $\tau(s)$  remains arbitrary!  $\nu(s)$  depends on  $\tau(s)$ . -> FAMILY of solutions.

## From fluid expansion to potential motion



The role of initial boundary conditions, EoS and freeze-out in hydro can be understood from potential motion!

# **Initial boundary conditions**

#### From the new family of exact solutions, the initial conditions:

#### **Initial coordinates:**

(nuclear geometry +
time of thermalization)

$$(X_0 Y_0 Z_0)$$

**Initial velocities:** 

$$(\dot{X}_{0}\,\dot{Y}_{0}\,\dot{Z}_{0})$$

(pre-equilibrium+ time of thermalization)

**Initial temperature:** 

 $T_0$ 

**Initial density:** 

 $n_0$ 

**Initial profile function:** 

 $\tau(s)$ 

(energy deposition and pre-equilibrium process)



## Role of initial temperature profile

- Initial temperature profile = arbitrary positive function
- **Infinitly rich class of solutions**
- Matching initial conditions for the density profile
  - T. Cs. Acta Phys. Polonica B37 (2006) 1001, hep-ph/0111139

$$\nu(s) = \frac{1}{\mathcal{T}(s)} \exp\left(-\frac{T_i}{2T_0} \int_0^s \frac{du}{\mathcal{T}(u)}\right)$$

**Homogeneous temperature** ⇒ **Gaussian density** 

$$\nu(s) = \exp(-s/2), \quad \mathcal{T}(s) = 1.$$
  $s = \frac{r_x^2}{X^2} + \frac{r_y^2}{Y^2} + \frac{r_z^2}{Z^2}$ 

$$s = \frac{r_x^2}{X^2} + \frac{r_y^2}{Y^2} + \frac{r_z^2}{Z^2}$$

Zimányi-Bondorf-Garpman profile:

**Buda-Lund profile:** 

$$\mathcal{T}(s) = \frac{1}{1+bs}$$

$$\nu(s) = (1+bs) \exp\left[-\frac{T_i}{2T_0}(s+bs^2/2)\right]$$

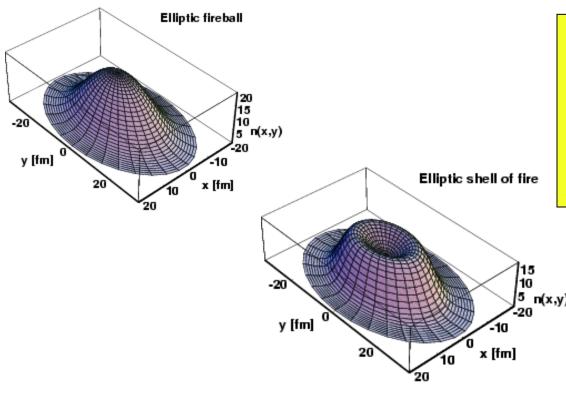
$$\mathcal{T}(s) = (1-s)\Theta(1-s)$$

$$\nu(s) = (1-s)^{\alpha}\Theta(1-s)$$

$$\mathcal{T}(s) = (1-s)\Theta(1-s)$$

$$\nu(s) = (1-s)^{\alpha}\Theta(1-s)$$

## Illustrated initial T-> density profiles

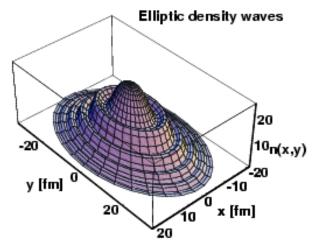


Determines density profile! Examples of density profiles

- Fireball
- Ring of fire
- Embedded shells of fire Exact integrals of hydro Scales expand in time

Time evolution of the scales (X,Y,Z) follows a classic potential motion. Scales at freeze out -> observables. info on history LOST!

No go theorem - constraints on initial conditions (penetrating probels) indispensable.



## Final (freeze-out) boundary conditions

From the new exact hydro solutions, the conditions to stop the evolution:

**Freeze-out temperature:** 

 $T_f$ 

**Final coordinates:** 

$$(X_f Y_f Z_f)$$

(cancel from measurables, diverge)

**Final velocities:** 

$$(\dot{X}_f \dot{Y}_f \dot{Z}_f)$$

(determine observables, tend to constants)

**Final density:** 

 $n_f$ 

(cancels from measurables, tends to 0)

**Final profile function:** 

 $\tau(s)$ 

(= initial profile function! from solution)



## **Role of the Equation of States:**

#### The potential depends

on  $\kappa = \delta \varepsilon / \delta p$ :

$$T_0 \left(\frac{V_0}{V}\right)^{1/\kappa}$$



Time evolution of the scales (X,Y,Z) follows a classic potential motion. Scales at freeze out determine the observables. Info on history <u>LOST!</u>
No go theorem - constraints on initial conditions
(information on spectra, elliptic flow of penetrating probels) indispensable.

The arrow hits the target, but can one determine g from this information??

## **Initial conditions <-> Freeze-out conditions:**

Different initial conditions

but

same freeze-out conditions

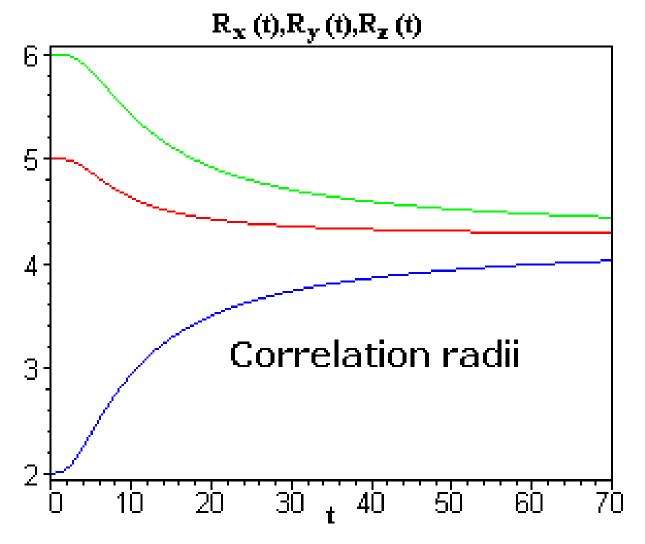
ambiguity!

Penetrating probes radiate through the time evolution!

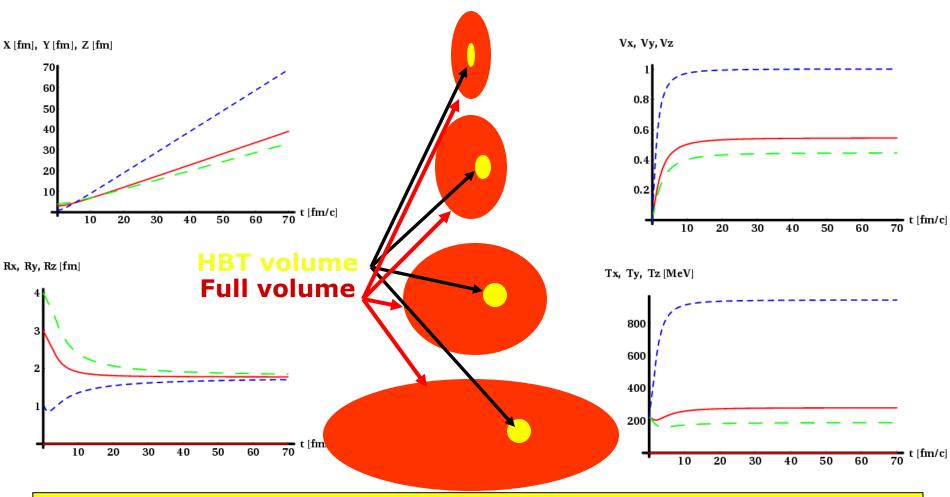


## Illustrations of exact hydro results

Propagate the hydro solution in time numerically:

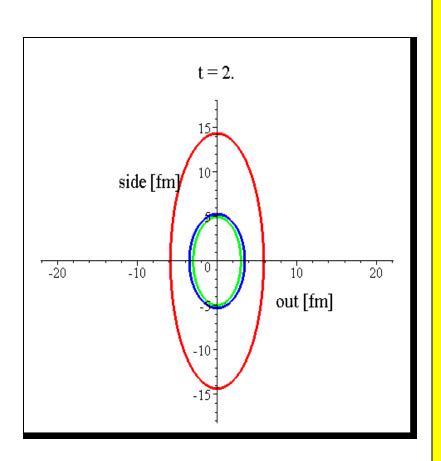


# Solution of the "HBT puzzle"



Geometrical sizes keep on increasing. Expansion velocities tend to constants. HBT radii  $R_x$ ,  $R_y$ ,  $R_z$  approach a direction independent constant. Slope parameters tend to direction dependent constants. General property, independent of initial conditions - a beautiful exact result.

## Geometrical & thermal & HBT radii



3d analytic hydro: exact time evolution

geometrical size (fugacity ~ const)

Thermal sizes (velocity ~ const)

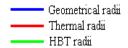
HBT sizes (phase-space density ~ const)

HBT dominated by the smaller of the geometrical and thermal scales

nucl-th/9408022, hep-ph/9409327 hep-ph/9509213, hep-ph/9503494

HBT radii approach a constant of time
HBT volume becomes spherical
HBT radii -> thermal ~ constant sizes

hep-ph/0108067, nucl-th/0206051 animation by Máté Csanád



# Scaling predictions of fluid dynamics

$$T'_x = T_f + m\dot{X}_f^2 ,$$
  

$$T'_y = T_f + m\dot{Y}_f^2 ,$$
  

$$T'_z = T_f + m\dot{Z}_f^2 .$$

- Slope parameters increase linearly with mass
- Elliptic flow is a universal function its variable w is proportional to transverse kinetic energy and depends on slope differences.

$$v_2 = \frac{I_1(w)}{I_0(w)}$$

$$v_2 = \frac{I_1(w)}{I_0(w)} \quad w = \frac{k_t^2}{4m} \left( \frac{1}{T_y'} - \frac{1}{T_x} \right), \quad w = \frac{E_K}{2T_*} \varepsilon$$

$$w = \frac{E_K}{2T_*}\varepsilon$$

Inverse of the HBT radii increase linearly with mass analysis shows that they are asymptotically the same

Relativistic correction: m -> m<sub>+</sub>

hep-ph/0108067, nucl-th/0206051

$$R_x'^{-2} = X_f^{-2} \left( 1 + \frac{m}{T_f} \dot{X}_f^2 \right),$$

$$R_y'^{-2} = Y_f^{-2} \left( 1 + \frac{m}{T_f} \dot{Y}_f^2 \right),$$

$$R_z'^{-2} = Z_f^{-2} \left( 1 + \frac{m}{T_f} \dot{Z}_f^2 \right).$$

## **Relativistic Perfect Fluids**

#### Rel. hydrodynamics of perfect fluids is defined by:

$$\partial_{\mu} (nu^{\mu}) = 0$$
$$\partial_{\mu} T^{\mu\nu} = 0$$

$$T^{\mu\nu} = (\varepsilon + p) u^{\mu}u^{\nu} - pg^{\mu\nu}$$

#### A recent family of exact solutions: nucl-th/0306004

$$u^{\mu} = \frac{x^{\mu}}{\tau}$$

$$n(t, \mathbf{r}) = n_0 \left(\frac{\tau_0}{\tau}\right)^3 \mathcal{V}(s)$$

$$p(t, \mathbf{r}) = p_0 \left(\frac{\tau_0}{\tau}\right)^{3+3/\kappa}$$

$$T(t, \mathbf{r}) = T_0 \left(\frac{\tau_0}{\tau}\right)^{3/\kappa} \frac{1}{\mathcal{V}(s)}$$

$$s = \frac{r_x^2}{X^2} + \frac{r_y^2}{Y^2} + \frac{r_z^2}{Z^2}, \qquad \rho = nT.$$

$$u_{\nu}u^{\mu}\partial_{\mu}p + (\epsilon + p)u^{\mu}\partial_{\mu}u_{\nu} - \partial_{\nu}p = 0,$$
  
$$u^{\mu}\partial_{\mu}T + \frac{1}{\kappa}T\partial_{\mu}u^{\mu} = 0.$$

$$s = \frac{r_x^2}{X^2} + \frac{r_y^2}{Y^2} + \frac{r_z^2}{Z^2} \,,$$

$$\epsilon = mn + \kappa p,$$
  
$$p = nT.$$

#### Overcomes two shortcomings of Bjorken's solution:

Yields finite rapidity distribution, includes transverse flow

**Hubble flow** ⇒ **lack of acceleration:** 

$$u^{\mu}\partial_{\mu}u_{\nu}=0$$

Accelerating, new rel. hydro solutions: nucl-th/0605070

## **Solutions of Relativistic Perfect Fluids**

#### A new family of exact solutions:

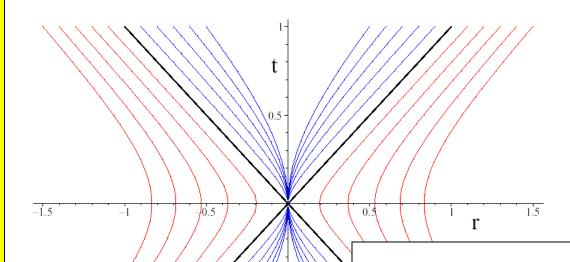
T. Cs, M. I. Nagy, M. Csanád: nucl-th/0605070

#### Overcomes two shortcomings of Bjorken's solution:

Finite Rapidity distribution ~ Landau's solution

**Includes relativistic acceleration** 

in 1+1 and 1+3 spherically symmetric



$$v = \tanh \lambda \eta,$$

$$n = n_0 \left(\frac{\tau_0}{\tau}\right)^{\lambda} \nu(s),$$

$$T = T_0 \left(\frac{\tau_0}{\tau}\right)^{\lambda} \frac{1}{\nu(s)}.$$

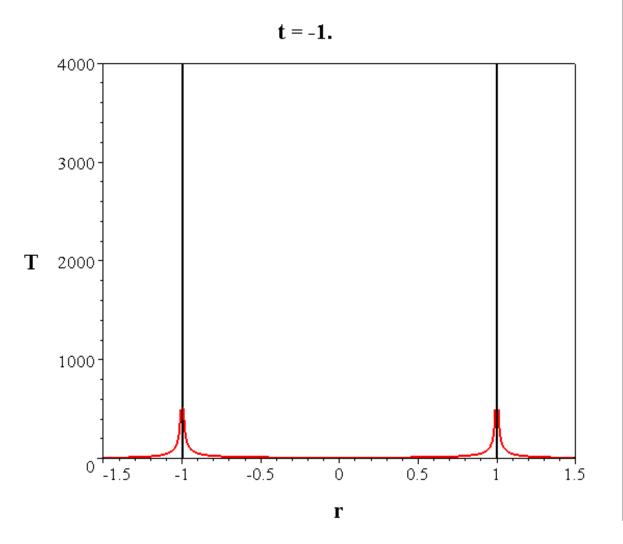
$$\frac{\mathrm{d}s}{\mathrm{d}t} = 0$$
.

$$|r| < |t| : s(\tau, \eta) = \left(\frac{\tau_0}{\tau}\right)^{\lambda - 1} \sinh\left((\lambda - 1)\eta\right),$$
  
$$|r| > |t| : s(\tau, \eta) = \left(\frac{\tau_0}{\tau}\right)^{\lambda - 1} \cosh\left((\lambda - 1)\eta\right).$$

## Animation of the new exact solution

nucl-th/0605070
dimensionless

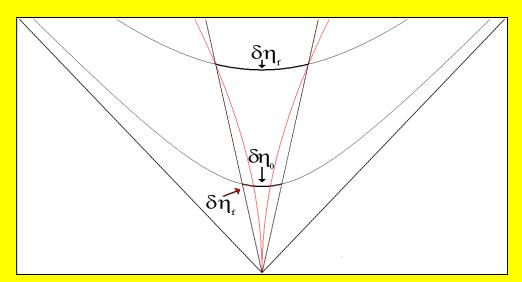
λ = 2
1+1 d
both internal
and external
looks like A+A



# nucl-th/0605070: advanced estimate of $\varepsilon_0$

#### Width of dn/dy distribution is due to acceleration:

acceleration yields longitudinal explosion, thus Bjorken estimate underestimates initial energy density by 50 %:



$$v = \tanh \lambda \eta,$$

$$n = n_0 \left(\frac{\tau_0}{\tau}\right)^{\lambda} \nu(s),$$

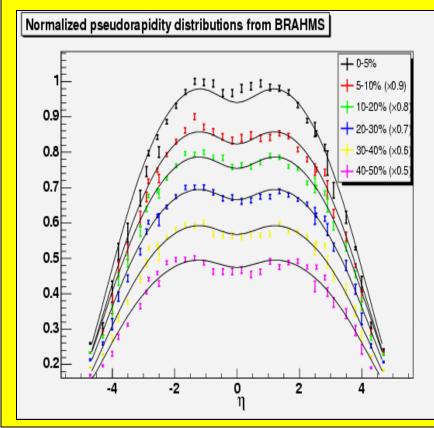
$$T = T_0 \left(\frac{\tau_0}{\tau}\right)^{\lambda} \frac{1}{\nu(s)}.$$

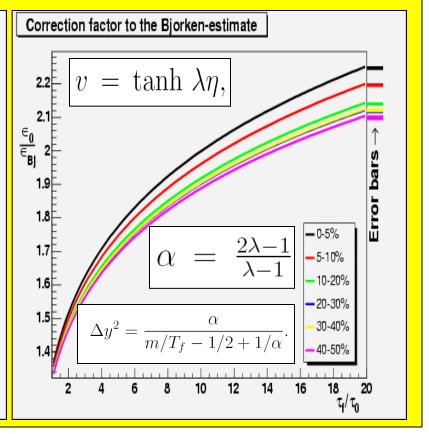
$$\epsilon_0 = \frac{\left\langle m_t \right\rangle}{R^2 \pi \tau_0} \frac{dn}{d\eta_0} = \epsilon_{Bj} \frac{dy}{d\eta_f} \frac{d\eta_f}{d\eta_0}$$

$$\epsilon_{0} = \frac{\left\langle m_{t} \right\rangle}{R^{2} \pi \tau_{0}} \frac{dn}{d\eta_{0}} = \epsilon_{Bj} \frac{dy}{d\eta_{f}} \frac{d\eta_{f}}{d\eta_{0}} = \frac{\epsilon_{0}}{\epsilon_{Bj}} \frac{\alpha}{\alpha - 2} \left(\frac{\tau_{f}}{\tau_{0}}\right)^{\frac{1}{\alpha - 2}} = (2\lambda - 1) \left(\frac{\tau_{f}}{\tau_{0}}\right)^{\lambda - 1}$$

# nucl-th/0605070: advanced estimate of $\varepsilon_0$

M. Csanád-> fits to BRAHMS dn/d $\eta$  data dn/d $\eta$  widths yields correction factors of  $\sim 2.0$  - 2.2 Yields inital energy density of  $\epsilon_0 \sim 10$ - 30 GeV/fm^3 a correction of  $\epsilon_0/\epsilon_{Bj} \sim 2$  as compared to PHENIX White Paper!



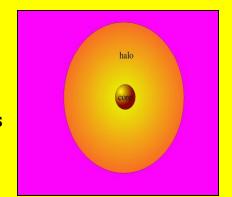


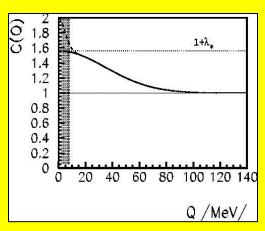
## **Principles for Buda-Lund hydro model**

- Analytic expressions for all the observables
- 3d expansion, local thermal equilibrium, symmetry
- Goes back to known exact hydro solutions:
  - nonrel, Bjorken, and Hubble limits, 1+3 d ellipsoids
  - but phenomenology, extrapolation for unsolved cases



- Core: perfect fluid dynamical evolution
- Halo: decay products of long-lived resonances
- Missing links: phenomenology needed
  - search for accelerating ellipsoidal rel. solutions
  - first accelerating rel. solution: nucl-th/0605070





## A useful analogy

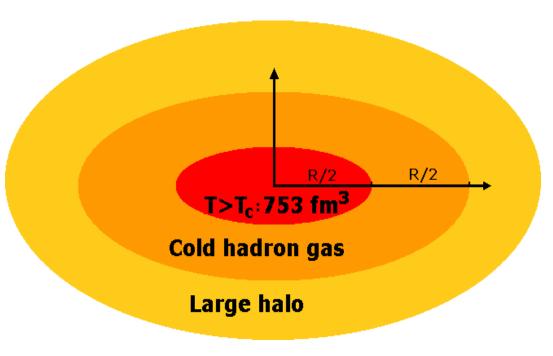
## Fireball at RHIC ⇔ our Sun

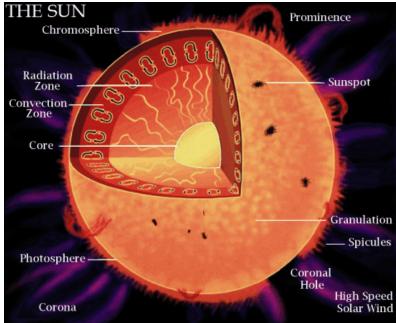
- Core
- Halo
- T<sub>0,RHIC</sub> ~ 210 MeV
- T<sub>surface,RHIC</sub> ~ 100 MeV



- **⇔** Solar wind
- $\Leftrightarrow$  T<sub>0,SUN</sub> ~ 16 million







## **Buda-Lund hydro model**

#### The general form of the emission function:

$$S_c(x, p)d^4x = \frac{g}{(2\pi)^3} \frac{p^{\mu}d^4\Sigma_{\mu}(x)}{\exp\left(\frac{p^{\nu}u_{\nu}(x)}{T(x)} - \frac{\mu(x)}{T(x)}\right) + s_q}$$

#### Calculation of observables with core-halo correction:

$$N_1(p) = \frac{1}{\sqrt{\lambda_*}} \int d^4x S_c(p, x)$$

$$C(Q, p) = 1 + \left| \frac{\tilde{S}(Q, p)}{\tilde{S}(0, p)} \right|^2 = 1 + \lambda_* \left| \frac{\tilde{S}_c(Q, p)}{\tilde{S}_c(0, p)} \right|^2$$

Assuming profiles for

flux, temperature, chemical potential and flow

## The generalized Buda-Lund model

The original model was for axial symmetry only, central coll.

In its general hydrodynamical form:

Based on 3d relativistic and non-rel solutions of perfect fluid dynamics:

$$S_c(x, p)d^4x = \frac{g}{(2\pi)^3} \frac{p^{\mu} d^4 \Sigma_{\mu}(x)}{\exp\left(\frac{p^{\nu} u_{\nu}(x)}{T(x)} - \frac{\mu(x)}{T(x)}\right) + s_q}$$

Have to assume special shapes:

**Generalized Cooper-Frye prefactor:** 

$$p^{\mu}d^{4}\Sigma_{\mu}(x) = p^{\mu}u_{\mu}(x)H(\tau)d^{4}x$$

$$H(\tau) = \frac{1}{(2\pi\Delta\tau^2)^{1/2}} \exp\left(-\frac{(\tau - \tau_0)^2}{2\Delta\tau^2}\right)$$

**Four-velocity distribution:** 

$$u^{\mu} = (\gamma, \sinh \eta_x, \sinh \eta_y, \sinh \eta_z)$$

**Temperature:** 

$$\frac{1}{T(x)} = \frac{1}{T_0} \left( 1 + \frac{T_0 - T_s}{T_s} s \right) \left( 1 + \frac{T_0 - T_e}{T_e} \frac{(\tau - \tau_0)^2}{2\Delta \tau^2} \right)$$

**Fugacity:** 

$$\frac{\mu(x)}{T(x)} = \frac{\mu_0}{T_0} - s$$

$$\frac{\mu(x)}{T(x)} = \frac{\mu_0}{T_0} - s$$

$$s = \frac{r_x^2}{X^2} + \frac{r_y^2}{Y^2} + \frac{r_z^2}{Z^2}$$

# Reminder: scaling laws, non-rel hydro

$$T'_x = T_f + m\dot{X}_f^2 ,$$
  

$$T'_y = T_f + m\dot{Y}_f^2 ,$$
  

$$T'_z = T_f + m\dot{Z}_f^2 .$$

- Slope parameters increase linearly with mass
- Elliptic flow is a <u>universal function</u> and variable w is proportional to transverse kinetic energy and depends on slope differences.

$$v_2 = \frac{I_1(w)}{I_0(w)}$$

$$v_2 = \frac{I_1(w)}{I_0(w)} \quad w = \frac{k_t^2}{4m} \left( \frac{1}{T_y'} - \frac{1}{T_x} \right), \quad w = \frac{E_K}{2T_*} \varepsilon$$

$$w = \frac{E_K}{2T_*}\varepsilon$$

Inverse of the HBT radii increase linearly with mass analysis shows that they are asymptotically the same

Relativistic correction: m -> m<sub>t</sub>

hep-ph/0108067, nucl-th/0206051

$$R_x'^{-2} = X_f^{-2} \left( 1 + \frac{m}{T_f} \dot{X}_f^2 \right),$$

$$R_y'^{-2} = Y_f^{-2} \left( 1 + \frac{m}{T_f} \dot{Y}_f^2 \right),$$

$$R_z'^{-2} = Z_f^{-2} \left( 1 + \frac{m}{T_f} \dot{Z}_f^2 \right).$$

# Scaling predictions: Buda-Lund hydro

$$T_x = T_0 + \overline{m}_t \dot{X}^2 \frac{T_0}{T_0 + \overline{m}_t a^2},$$

$$\overline{m}_t = m_t \cosh(\eta_s - y).$$

- Slope parameters increase linearly with transverse mass
- Elliptic flow is same universal function.
- Scaling variable w is prop. to generalized transv. kinetic energy and depends on effective slope diffs.

$$v_2 = \frac{I_1(w)}{I_0(w)} \bigg| \quad w = \frac{E_K}{2T_*} \varepsilon$$

$$w = \frac{E_K}{2T_*}\varepsilon$$

$$\varepsilon = \frac{T_x - T_y}{T_x + T_y}.$$

$$E_K = \frac{p_t^2}{2\overline{m}_t}$$

$$E_K = \frac{p_t^2}{2\overline{m}_t} \qquad \boxed{ \frac{1}{T_*} = \frac{1}{2} \left( \frac{1}{T_x} + \frac{1}{T_y} \right). }$$

Inverse of the HBT radii increase linearly with mass analysis shows that they are asymptotically the same

Relativistic correction: m -> m<sub>+</sub>

hep-ph/0108067, nucl-th/0206051

$$\frac{1}{R_{i,i}^2} = \frac{B(x_s, p)}{B(x_s, p) + s_q} \left( \frac{1}{X_i^2} + \frac{1}{R_{T,i}^2} \right)$$

$$\frac{1}{R_{T,i}^2} = \frac{m_t}{T_0} \left( \frac{a^2}{X_i^2} + \frac{\dot{X}_i^2}{X_i^2} \right)$$

## Some analytic Buda-Lund results

#### **HBT radii widths:**

$$\frac{1}{R_{i,i}^2} = \frac{B(x_s, p)}{B(x_s, p) + s_q} \left( \frac{1}{X_i^2} + \frac{1}{R_{T,i}^2} \right) = \frac{1}{R_{T,i}^2} = \frac{m_t}{T_0} \left( \frac{a^2}{X_i^2} + \frac{\dot{X}_i^2}{X_i^2} \right)$$

$$\frac{1}{R_{T,i}^2} = \frac{m_t}{T_0} \left( \frac{a^2}{X_i^2} + \frac{\dot{X}_i^2}{X_i^2} \right)$$

 $a^2 = \frac{T_0 - T_s}{T_s} = \left\langle \frac{\Delta T}{T} \right\rangle$ 

#### Slopes, effective temperatures:

$$\overline{m}_t = m_t \cosh(\eta_s - y).$$

$$\frac{1}{T_*} = \frac{1}{2} \left( \frac{1}{T_x} + \frac{1}{T_y} \right). \qquad T_x = T_0 + \overline{m}_t \dot{X}^2 \frac{T_0}{T_0 + \overline{m}_t a^2},$$

$$T_x = T_0 + \overline{m}_t \, \dot{X}^2 \frac{T_0}{T_0 + \overline{m}_t a^2},$$

#### Flow coefficients are <u>universal</u>:

$$v_{2n} = \frac{I_n(w)}{I_0(w)}$$

$$v_{2n+1} = 0$$

$$w = \frac{E_K}{2T_*} \varepsilon$$

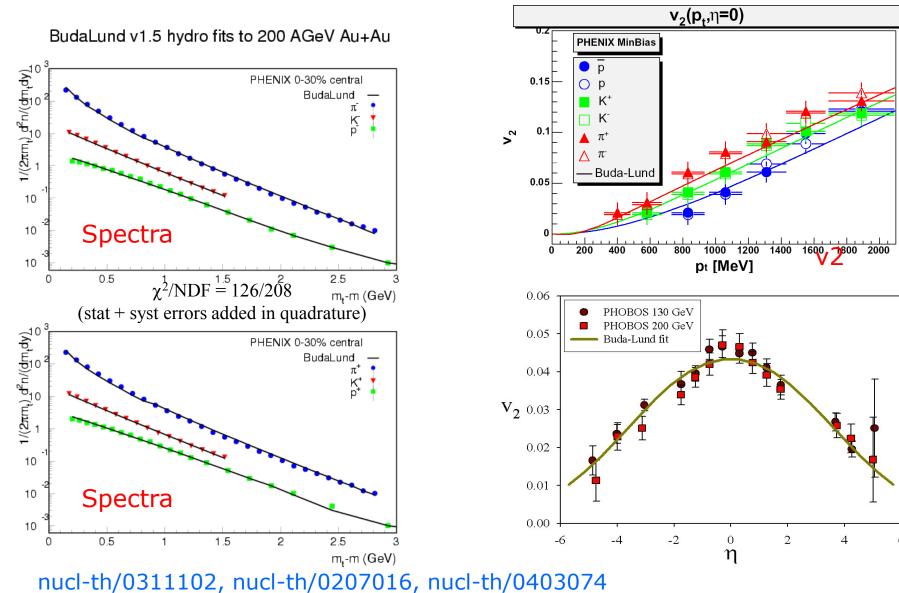
$$E_K = \frac{p_t^2}{2\overline{m}_t}$$

$$w = \frac{E_K}{2T_*}\varepsilon$$

$$\varepsilon = \frac{T_x - T_y}{T_x + T_y}.$$

$$E_K = \frac{p_t^2}{2\overline{m}_t}$$

# **Buda-Lund hydro and Au+Au@RHIC**



35 T. Csörgő @ PHENIX Focus, BNL, 2006/6/13

## Femptoscopy signal of sudden hadronization

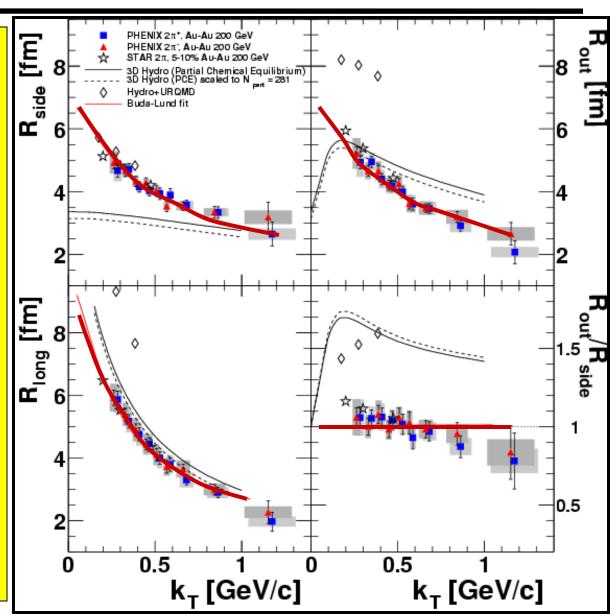
Buda-Lund hydro fit indicates hydro predicted (1994-96) scaling of HBT radii

T. Cs, L.P. Csernai hep-ph/9406365

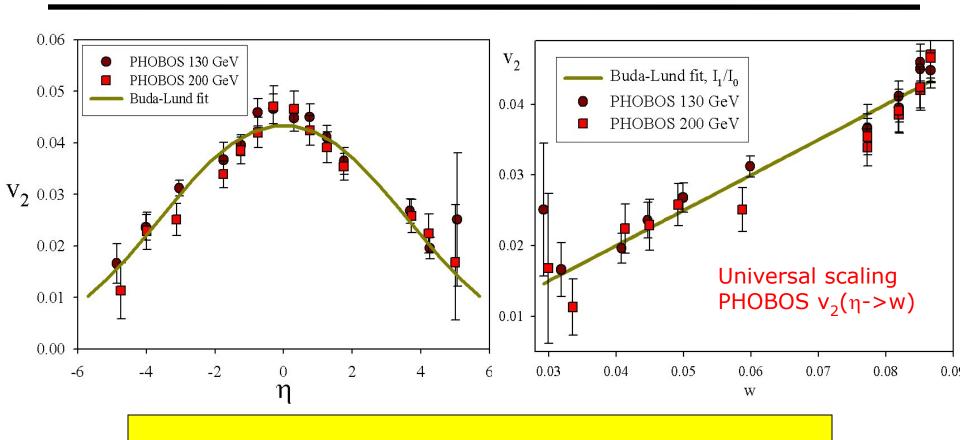
T. Cs, B. Lörstad hep-ph/9509213

Hadrons with T>T<sub>c</sub>:
a hint for
cross-over

M. Csanád, T. Cs, B. Lörstad and A. Ster, nucl-th/0403074



#### **Confirmation**



see nucl-th/0310040 and nucl-th/0403074, R. Lacey@QM2005/ISMD 2005 A. Ster @ QM2005.

## Hydro scaling of slope parameters

#### **Buda-Lund hydro prediction:**

#### Exact non-rel. hydro solution:

$$T_{*,i} = T_0 + m_t \, \dot{X}_i^2 \frac{T_0}{T_0 + m_t a^2}$$

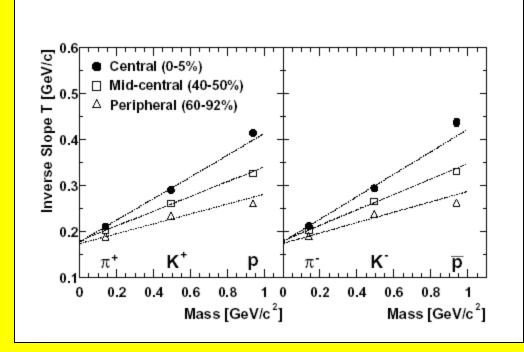


$$T'_x = T_f + m\dot{X}_f^2 ,$$
  

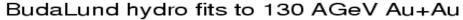
$$T'_y = T_f + m\dot{Y}_f^2 ,$$
  

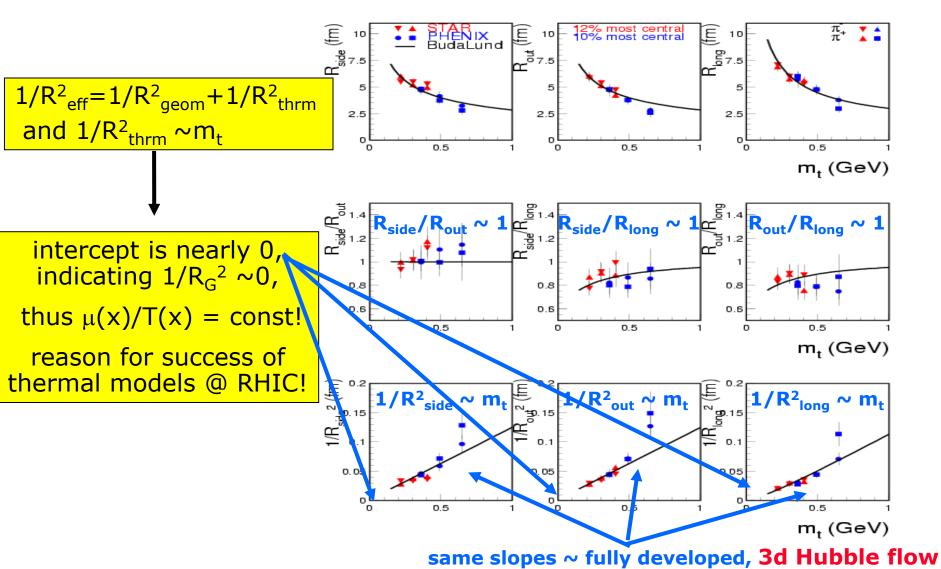
$$T'_z = T_f + m\dot{Z}_f^2 .$$

#### **PHENIX data:**



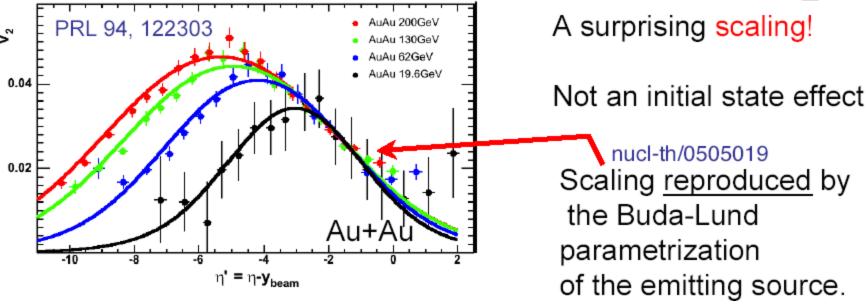
#### Hydro scaling of Bose-Einstein/HBT radii





### Hydro scaling of elliptic flow

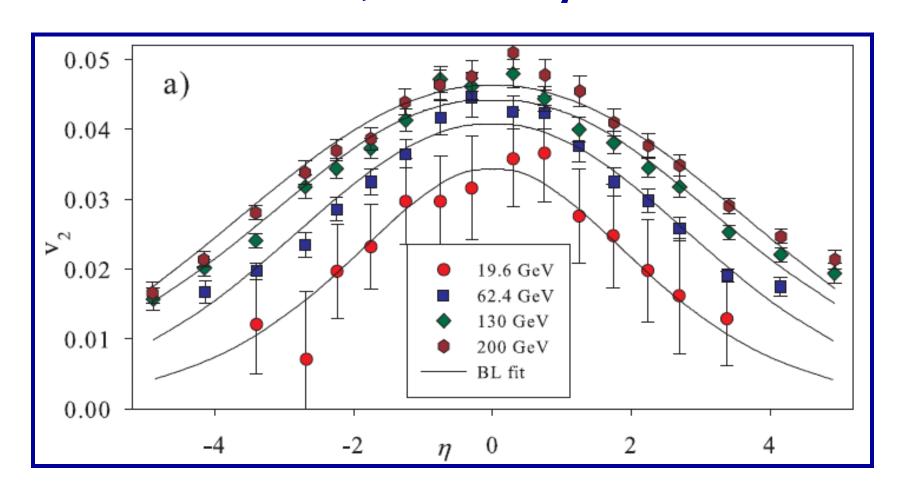
Extended longitudinal scaling: v<sub>2</sub>



G. Veres, PHOBOS data, proc QM2005 Nucl. Phys. A774 (2006) in press

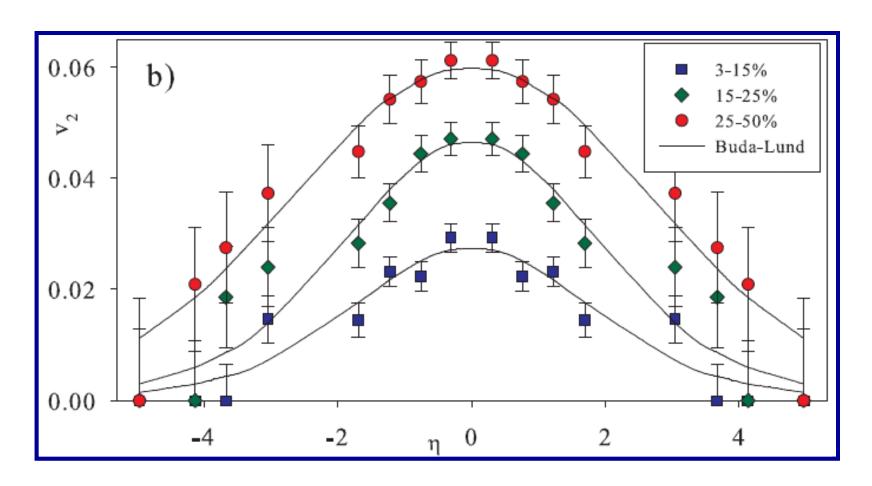
# Hydro scaling of $v_2$ and $\sqrt{s}$ dependence

# PHOBOS, nucl-ex/0406021



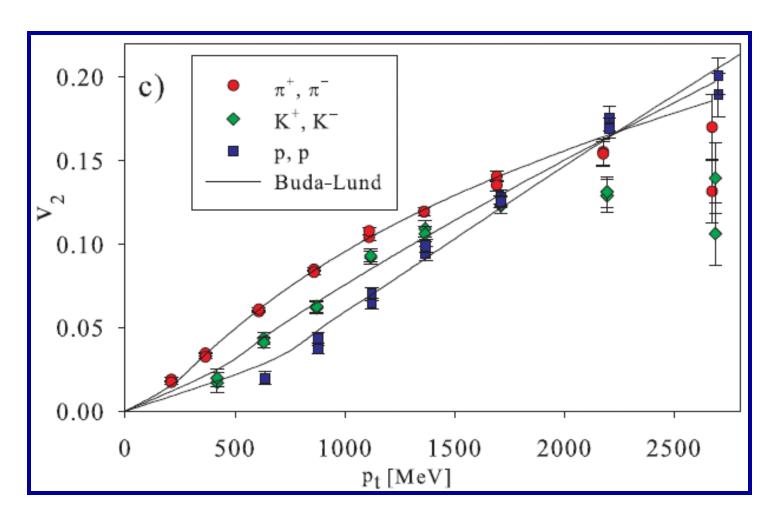
#### Universal scaling and $v_2$ (centrality, $\eta$ )

## PHOBOS, nucl-ex/0407012



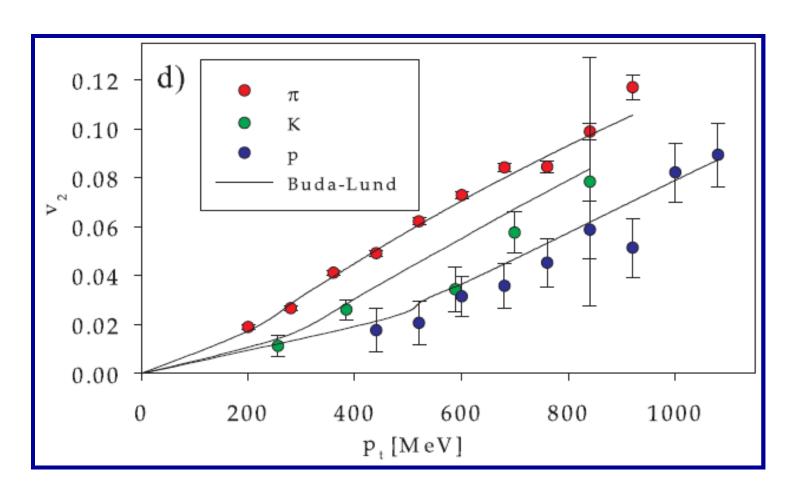
#### Universal v2 scaling and PID dependence

## PHENIX, nucl-ex/0305013

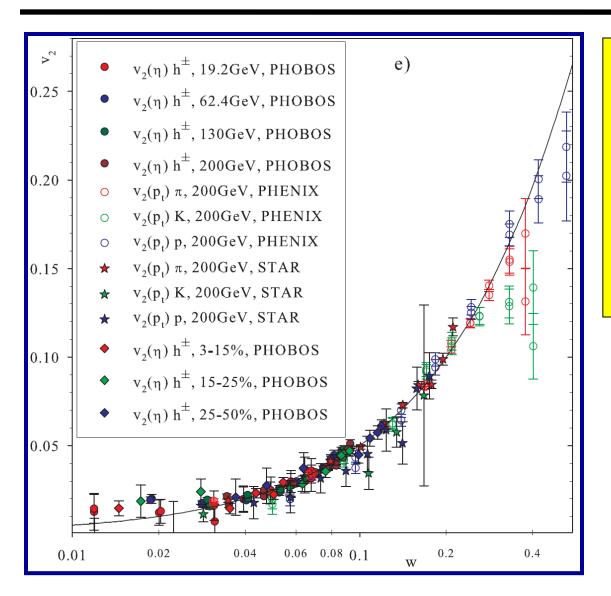


#### Universal scaling and fine structure of v2

# STAR, nucl-ex/0409033



#### Summary on universal hydro scaling of v<sub>2</sub>

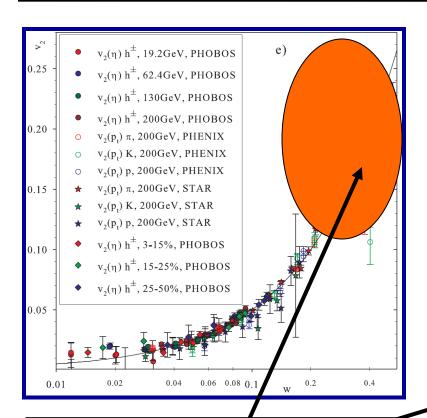


Black line:
Theoretically
predicted, universal
scaling function
from analytic works
on perfect fluid
hydrodynamics:

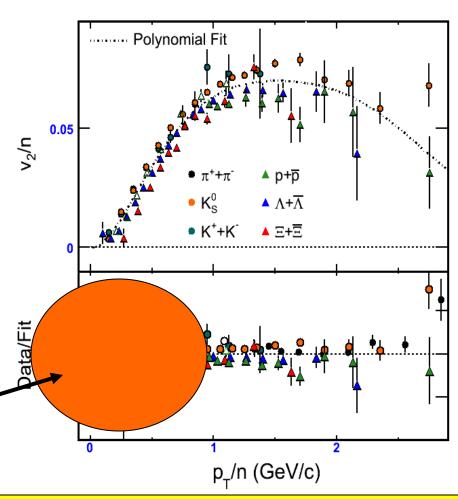
$$v_2 = \frac{I_1(w)}{I_0(w)}$$

hep-ph/0108067, nucl-th/0310040

## Scaling and scaling violations



Universal hydro scaling breaks where scaling with number of VALENCE QUARKS sets in, p<sub>t</sub> ~ 1-2 GeV Fluid of QUARKS!!



R. Lacey and M. Oldenburg, proc. QM'05 A. Taranenko et al, PHENIX PPG062: nucl-ex/0608033

## Understanding hydro results

**New exact solutions of 3d nonrelativistic hydrodynamics:** Hydro problem equivalent to potential motion (a shot)! **Shot of an arrow: Hydro: Desription of data Hitting the target Initial** co elocity **Equation** ntial Freeze-d Data con bout the potentia **Differen** exactly t aneously (!) **EoS** and co-varied with the potential Universal scaling of v<sub>2</sub> In a perfect shot, the shape of trajectory is a parabola Drag force of air Viscosity effects

## **Summary**

Au+Au elliptic flow data at RHIC satisfy the UNIVERSAL scaling laws predicted (2001, 2003) by the (Buda-Lund) hydro model, based on exact solutions of PERFECT FLUID hydrodynamics

quantitative evidence for a perfect fluid in Au+Au at RHIC

scaling breaks in  $p_t$  at  $\sim 1.5$  GeV, in rapidity at  $\sim |y| > y_{may}$  - 0.5 Search for establishing the domain of applicability started. Scaling of HBT radii and spectra: first tests passed, further tests going on.